

maxima: orbit equation , version 1.0
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General equation : (orbit or trajectory) of space vehicle
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part 1 : Energy

step 1:kinetic energy:T

from mechanics : $T=1/2*m*V^2$, kinetic energy from motion.

p1a: $T=1/2*m*V^2$;

$$T = \frac{V^2 m}{2}$$

m: small mass

V: velocity of small body around the large mass

a) polair equation for V, velocity V [m/s]

p1a1: $V^2=(\text{'diff(r,t)})^2+(r*\text{'diff(theta,t)})^2$;

$$V^2 = r^2 \left(\frac{d}{dt} \theta \right)^2 + \left(\frac{d}{dt} r \right)^2$$

r: distance between large mass and small mass

b) final kinetic : equation

p1a2:subst([p1a1],p1a);

$$T = \frac{m \left(r^2 \left(\frac{d}{dt} \theta \right)^2 + \left(\frac{d}{dt} r \right)^2 \right)}{2}$$

p1a21:expand(p1a2);

$$T = \frac{m r^2 \left(\frac{d}{dt} \theta \right)^2}{2} + \frac{m \left(\frac{d}{dt} r \right)^2}{2}$$

step 2: pontential energy:P

from mechanics : P

a) (F)orce : motion small mass:m around gravitation field of large mass :M

p1b:F=G*(M*m)/r^2; [N : kg*m/s^2];

$$F = \frac{G M m}{r^2} \quad \left[\frac{\text{kg m}}{\text{s}^2} \right]$$

G: universal gravitational constant:6.67*10⁻¹¹ m^3/(kg)*(s)^2

M: mass large mass , earth → 5.98*10²⁴ kg

m: small mass

p1ba1:Γ=G*M;

$$\Gamma = G M$$

p1ba2:reverse(p1ba1);

$$G M = \Gamma$$

b) P : integration of (F)orce :from infinity to r

assume(r>0);

$$[r > 0]$$

p1b1:"integrate(1,P,0,P)=integrate(rhs(p1b),r,∞,r);

$$P = - \left(G M m \int_r^{\infty} \frac{1}{r^2} dr \right)$$

p1b1a:lratsubst([reverse(p1ba1)],ev(p1b1,nouns));

$$P = - \left(\frac{m \Gamma}{r} \right)$$

step 3: E ,total Energy

p1c:E= T+P;

$$E = T + P$$

use : p1a2 & p1b1a

a) Energy :E [Joule =J=N*m]

p1c1:subst([p1a2,p1b1a],p1c);

$$E = \frac{m \left(r^2 \left(\frac{d}{dt} \theta \right)^2 + \left(\frac{d}{dt} r \right)^2 \right)}{2} - \frac{m \Gamma}{r}$$

b) Energy per unit mass : $E_m = E/m$

p1c2: $E_m = \text{expand}(\text{rhs}(\text{p1c1})/m);$

$$E_m = \frac{r^2 \left(\frac{d}{dt} \theta \right)^2}{2} - \frac{\Gamma}{r} + \frac{\left(\frac{d}{dt} r \right)^2}{2}$$

part 2: (La)grangian & (La)grangian (equation)

→ (La)grangian

p2a: $La = T - P;$

p2a1: $La = \text{subst}([p1a2, p1b1a], \text{rhs}(p2a));$

$$La = T - P \quad La = \frac{m \left(r^2 \left(\frac{d}{dt} \theta \right)^2 + \left(\frac{d}{dt} r \right)^2 \right)}{2} + \frac{m \Gamma}{r}$$

→ (La)grangian (equation)

depends $([\theta, r], t);$

$[\theta(t), r(t)]$

p2b: $\text{Laeq}(va, B) := \text{diff}(\text{diff}(B, \text{diff}(va, t)), t) - \text{diff}(B, va) = 0;$

$$\text{Laeq}(va, B) := \frac{d}{dt} \left(\frac{d}{dt} \left(\frac{d}{dt} B \right) \right) - \frac{d}{dva} B = 0$$

→ $va: \theta, \text{Laeq}(La, va);$

p2b1a: $\text{Laeq}(\theta, \text{rhs}(p2a1));$

$$m r^2 \left(\frac{d^2}{dt^2} \theta \right) + 2 m r \left(\frac{d}{dt} r \right) \left(\frac{d}{dt} \theta \right) = 0$$

p2b2a: $'\text{diff}((m \cdot r^2 \cdot \text{diff}(\theta, t)), t) = \text{diff}((m \cdot r^2 \cdot \text{diff}(\theta, t)), t);$

$$\frac{d}{dt} \left(m r^2 \left(\frac{d}{dt} \theta \right) \right) = m r^2 \left(\frac{d^2}{dt^2} \theta \right) + 2 m r \left(\frac{d}{dt} r \right) \left(\frac{d}{dt} \theta \right)$$

final result : $va = \theta$

p2b2b: $\text{ratsubst}([p2b1a], p2b2a);$

$$\frac{d}{dt} \left(m r^2 \left(\frac{d}{dt} \theta \right) \right) = 0$$

rule : $d(A)/dt = 0, A = \text{constant} \Rightarrow A = \text{angular momentum.}$

p2b2c: $\text{part}(\text{lhs}(p2b2b), 1) = A;$

$$m r^2 \left(\frac{d}{dt} \theta \right) = A$$

→ $va: r, \text{Laeq}(La, va);$

p2c1: $\text{Laeq}(r, \text{rhs}(p2a1));$

$$-\left(m r \left(\frac{d}{dt} \theta \right)^2 \right) + \frac{m \Gamma}{r^2} + m \left(\frac{d^2}{dt^2} r \right) = 0$$

definition : angular momentum per unit mass

p2c2: $h = r^2 \cdot \text{diff}(\theta, t);$

$$h = r^2 \left(\frac{d}{dt} \theta \right)$$

→ use : $p2c1 \cdot r^3/m$

p2c3: $\text{expand}(p2c1 \cdot r^3/m);$

$$-\left(r^4 \left(\frac{d}{dt} \theta \right)^2 \right) + r \Gamma + r^3 \left(\frac{d^2}{dt^2} r \right) = 0$$

→ use : $p2c2$ into $p2c3 \Rightarrow */r^3$

p2c4: $\text{expand}(\text{ratsubst}([\text{reverse}(p2c2)], p2c3)/r^3);$

$$\frac{\Gamma}{r^2} + \frac{d^2}{dt^2} r - \frac{h^2}{r^3} = 0$$

$r = 1/u, h = \text{diff}(\theta, t)/u^2$

depends $([u], t);$

$[u(t)]$

< $dr/dt = ?$ >, first order differential from r with respect to t

→ $dr/dt = ?$, $\text{subst } r = 1/u$

p4d1: $\text{subst}([r = 1/u], \text{diff}(r, t, 1));$

$$\frac{d}{dt} \frac{1}{u}$$

p4d1:ev(p4d1,nouns);

$$-\left(\frac{\frac{d}{dt} u}{u^2}\right)$$

depends([u],θ);

$$[u(\theta)]$$

→ du/dt = ? ,subst du/dt = du/dθ*dθ/dt

p4d2:ratsubst([diff(u,t)=diff(u,θ)*diff(θ,t)],ev(p4d1,nouns));

$$-\left(\frac{\left(\frac{d}{d\theta} u\right)\left(\frac{d}{dt} \theta\right)}{u^2}\right)$$

→ use h , p2c2

p4d3:subst([r=1/u],p2c2);

$$h = \frac{\frac{d}{dt} \theta}{u^2}$$

→ subst p4d3 into p4d2

p4d4:ratsubst([reverse(p4d3)],p4d2);

$$-\left(h \left(\frac{d}{d\theta} u\right)\right)$$

< d(dr/dt)/dt = ?> , second order differential from r with respect to t

use :p4d4

p4e1:diff(r,t,2)=diff(p4d4,t);

$$\frac{d^2}{dt^2} r = -\left(h \left(\frac{d^2}{d\theta^2} u\right) \left(\frac{d}{dt} \theta\right)\right)$$

use : p4d3 ,solve : diff(θ,t)

p4e1a:solve(p4d3,diff(θ,t));

$$\left[\frac{d}{dt} \theta = h u^2\right]$$

p4e2:subst(%,p4e1);

$$\frac{d^2}{dt^2} r = -\left(h^2 u^2 \left(\frac{d^2}{d\theta^2} u\right)\right)$$

part 3: final differential equation computing orbit equation

use :p2c4

p5f1:p2c4;

$$\frac{\Gamma}{r^2} + \frac{d^2}{dt^2} r - \frac{h^2}{r^3} = 0$$

use : p4e2 into p2c4

p5f2:subst([p4e2],p5f1);

$$\frac{\Gamma}{r^2} - h^2 u^2 \left(\frac{d^2}{d\theta^2} u\right) - \frac{h^2}{r^3} = 0$$

p5f2a:coeff(lhs(p5f2),diff(u,θ,2));

$$-(h^2 u^2)$$

use :p5f2/p5f2a

p5f3:expand(p5f2/p5f2a);

$$-\left(\frac{\Gamma}{h^2 r^2 u}\right) + \frac{d^2}{d\theta^2} u + \frac{1}{r^3 u^2} = 0$$

use: p5f3 , subst r=1/u

p5f4:subst([r=1/u],p5f3);

$$-\left(\frac{\Gamma}{h^2}\right) + \frac{d^2}{d\theta^2} u + u = 0$$

p5f5:ode2(p5f4,u,θ);

$$u = \%k1 \sin(\theta) + \%k2 \cos(\theta) + \frac{\Gamma}{h^2}$$

part 4:find polair equation , u =1/r.

form : r=p/(1+e*cos(θ-C))

subst: u=1/r into p5f5

p6e1:subst([u=1/r],p5f5);

$$\frac{1}{r} = \%k1 \sin(\theta) + \%k2 \cos(\theta) + \frac{\Gamma}{h^2}$$

p6e1a:r=p/(1+e*cos(θ-C));/*ref for conic section*/

$$r = \frac{p}{e \cos(\theta - C) + 1}$$

e:eccentricity

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if e=0 the orbit is a circle

if e<1 the orbit is an ellipse

if e=1 the orbit is a parabola

if e> 1 the orbit is a hyperbola

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C = phase angle ,below pos x-axis, "ref : angles "

θ = true anomaly,

r=p ,as θ-BB =π/2

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solve p6e1 for r

p6e2:solve(p6e1,r);

$$\left[r = \frac{h^2}{\%k1 h^2 \sin(\theta) + \%k2 h^2 \cos(\theta) + \Gamma} \right]$$

differential eq : second order two constant we must find

use :numerator of r

p6e2a:num(rhs(p6e2[1]))/Γ; /*find form :p6e1a*/

$$\frac{h^2}{\Gamma}$$

use : denominator of r

p6e2aa:expand(denom(rhs(p6e2[1])));

$$\%k1 h^2 \sin(\theta) + \%k2 h^2 \cos(\theta) + \Gamma$$

use : {%k1,%k2} → {AA,BB}

p6e2a1:p6e2aa=AA·h^2·cos(θ-BB)+Γ; /*find form :p6e1a*/

$$\%k1 h^2 \sin(\theta) + \%k2 h^2 \cos(\theta) + \Gamma = AA h^2 \cos(\theta - BB) + \Gamma$$

p6e2a2:expand(p6e2a1/Γ); /*find form :p6e1a*/

$$\frac{\%k1 h^2 \sin(\theta)}{\Gamma} + \frac{\%k2 h^2 \cos(\theta)}{\Gamma} + 1 = \frac{AA h^2 \cos(\theta - BB)}{\Gamma} + 1$$

use : p6e2a/p6e2a2

p6e3: r =p6e2a/rhs(p6e2a2);

$$r = \frac{h^2}{\Gamma \left(\frac{AA h^2 \cos(\theta - BB)}{\Gamma} + 1 \right)}$$

1e) r=p , when true anomaly , θ-BB =π/2

p6e3a: subst([r=p, θ-BB =π/2],p6e3);

$$p = \frac{h^2}{\Gamma}$$

pe3a1:ratsubst([reverse(p6e3a)],p6e3);

$$r = \frac{p}{AA p \cos(\theta - BB) + 1}$$

see:p6e1a:r=p/(1+e*cos(θ-C));/*ref for conic section,see my website*/

2e) AA.p =e

pe3b:ratsubst([AA·p =e],pe3a1);

$$r = \frac{p}{e \cos(\theta - BB) + 1}$$

rem :

pe3c:(denom(rhs(pe3b))·r=p)^2;

$$r^2 (e \cos(\theta - BB) + 1)^2 = p^2$$

pe3c1:reverse(pe3c);

$$p^2 = r^2 (e \cos(\theta - BB) + 1)^2$$

part 5: How to find e , for orbit use E or Em

use : E = T+P , p1c ,conservation energy

a) T use "conic section : r"

use : compute velocity V

pd1:diff(pe3b,t);

$$\frac{d}{dt} r = \frac{e p \left(\frac{d}{dt} \theta \right) \sin(\theta - BB)}{(e \cos(\theta - BB) + 1)^2}$$

pd1a:part(rhs(p1c),1)=(subst([pd1],rhs(p1a21))));

$$T = \frac{e^2 m p^2 \left(\frac{d}{dt} \theta \right)^2 \sin^2(\theta - BB)}{2 (e \cos(\theta - BB) + 1)^4} + \frac{m r^2 \left(\frac{d}{dt} \theta \right)^2}{2}$$

use :pe3c1

pd1b:subst([pe3c1],pd1a);

$$T = \frac{e^2 m r^2 \left(\frac{d}{dt} \theta \right)^2 \sin^2(\theta - BB)}{2 (e \cos(\theta - BB) + 1)^2} + \frac{m r^2 \left(\frac{d}{dt} \theta \right)^2}{2}$$

use :p2c2

pdd: subst(solve(p2c2,diff(θ,t)),pd1b);

$$T = \frac{e^2 h^2 m \sin^2(\theta - BB)}{2 r^2 (e \cos(\theta - BB) + 1)^2} + \frac{h^2 m}{2 r^2}$$

use : pe3c

pd1c:factor(trigsimp(factor(lratsubst([pe3c],pdd))));

$$T = \frac{h^2 m (2 e \cos(\theta - BB) + e^2 + 1)}{2 p^2}$$

combination : factor(trigsimp(factor(?)))

use : trig function are inside

b) P use "conic section : r"

use :p1b1a

pd1d:p1b1a;

$$P = - \left(\frac{m \Gamma}{r} \right)$$

use :p6e1a

pd1da:subst([p6e1a],pd1d);

$$P = - \left(\frac{m \Gamma (e \cos(\theta - C) + 1)}{p} \right)$$

use:p1c1

pdf:subst([pd1da,pd1c],E=T+P);

$$E = \frac{h^2 m (2 e \cos(\theta - BB) + e^2 + 1)}{2 p^2} - \frac{m \Gamma (e \cos(\theta - C) + 1)}{p}$$

use:p6e3a

use : BB=C

pdf1:ratsimp(subst([p6e3a,BB=C],pdf));

$$E = \frac{(e^2 - 1) m \Gamma^2}{2 h^2}$$

pdf2:combine(expand(pdf1));

$$E = \frac{e^2 m \Gamma^2 - m \Gamma^2}{2 h^2}$$

pdf3:solve(pdf2,e);

$$\left[e = - \left(\frac{\sqrt{\Gamma^2 + \frac{2 E h^2}{m}}}{\Gamma} \right), e = \frac{\sqrt{\Gamma^2 + \frac{2 E h^2}{m}}}{\Gamma} \right]$$

use pdf3, because all values are pos

pdf4:pdf3[2];

$$e = \frac{\sqrt{\Gamma^2 + \frac{2 E h^2}{m}}}{\Gamma}$$

Orbits (= Trajectory of the small mass = m around a large mass =M)

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Orb: matrix(
[circle, e=0, Nan],
[Ellipse, e<1, 1/2*m*V^2<G*M*m/r],
[Parabola, e=1, 1/2*m*V^2=G*M*m/r],
[Hyperbola, e>1, 1/2*m*V^2>G*M*m/r]
);

```

$$\begin{pmatrix} \text{circle} & e=0 & \text{Nan} \\ \text{Ellipse} & e<1 & \frac{V^2 m}{2} < \frac{G M m}{r} \\ \text{Parabola} & e=1 & \frac{V^2 m}{2} = \frac{G M m}{r} \\ \text{Hyperbola} & e>1 & \frac{V^2 m}{2} > \frac{G M m}{r} \end{pmatrix}$$

```
pcf:E = 1/2*m*V^2-G*M/r;
```

$$E = \frac{V^2 m}{2} - \frac{m \Gamma}{r}$$

a) computation , e=0, circular velocity

```
use :pdf2
```

```
pcf1:subst([e=0],pdf2);/*circle,info from table*/
```

$$E = - \left(\frac{m \Gamma^2}{2 h^2} \right)$$

```
pcf2:expand(subst([pcf],pcf1)/m);
```

$$\frac{V^2}{2} - \frac{\Gamma}{r} = - \left(\frac{\Gamma^2}{2 h^2} \right)$$

```
use:pe3c1
```

```
pcf3:reverse(subst([e=0],pe3c1));
```

$$\frac{2}{r} = \rho^2$$

```
use :p6e3a
```

```
pcf4:reverse(p6e3a*Γ);
```

$$h^2 = \rho \Gamma$$

```
subst : pcf4 into pcf2,use pcf3
```

```
pcf5:subst([pcf4,p=r],pcf2);
```

$$\frac{V^2}{2} - \frac{\Gamma}{r} = - \left(\frac{\Gamma}{2 r} \right)$$

```
pcf6:solve(pcf5,V);
```

$$\left[V = - \left(\frac{\sqrt{\Gamma}}{\sqrt{r}} \right), V = \frac{\sqrt{\Gamma}}{\sqrt{r}} \right]$$

because both pos : e=0

```
print( "V:Circular Velocity =", (rhs(pcf6[2])))$
```

$$V:\text{Circular Velocity} = \frac{\sqrt{\Gamma}}{\sqrt{r}}$$

b) computation , e=1, parabolic velocity

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pbf:subst([E=0],pcf);
```

$$0 = \frac{V^2 m}{2} - \frac{m \Gamma}{r}$$

```
pbf1:solve(pbf,V);
```

$$\left[V = - \left(\frac{\sqrt{2} \sqrt{\Gamma}}{\sqrt{r}} \right), V = \frac{\sqrt{2} \sqrt{\Gamma}}{\sqrt{r}} \right]$$

```
print( "V:Parabolic Velocity =", (rhs(pbf1[2])))$
```

$$V:\text{Parabolic Velocity} = \frac{\sqrt{2} \sqrt{\Gamma}}{\sqrt{r}}$$

that is called escape velocity : r=re ,radius of earth.

example : computation , 0< e< 1,elliptic orbit

rocket launch (= space vehicle) burnout velocity t @ 8km/s above the earth , position of rocket 30 ° degree above earth equator at distance above earth 8000 km where rocket burnout ,velocity angle 10° above the horizontal horizon .

rem : horizontal horizon is perpendicular to burnout distance (= RE +height above the eart)

compute : the elements of the orbit :

$$r = p / (1 + e \cdot \cos(\theta - C))$$

solution :

1e) height above the earth [m]

$$r = R_e + h$$

Re: 6.44 · 10⁶ ; /*radius of the earth*/

rh : 8.0 · 10⁶ ; /*given*/

$$6440000.0 \quad 8000000.0$$

q1a: hh = rh - Re;

$$hh = 1560000.0$$

2e) $\Gamma = G \cdot M_e$

Γ : 3.986 · 10¹⁴ ; /*m³/s²*/

$$3.986 \cdot 10^{14}$$

3e) $h = r \cdot V_0 = r \cdot V \cdot \cos(C)$

V: 8000 ; /*m/s*/

C: float(7 · π / 180) ; /*rad*/

sa: h = rh · V · cos(C);

$$8000 \quad 0.12217304763960307 \quad h = 6.352295370504461 \cdot 10^{10}$$

4e) $p = h^2 / \Gamma$

p: ev(subst([%], h^2 / Γ), nouns);

$$1.0123345828934371 \cdot 10^7$$

5e) $e = \sqrt{1 + 2 \cdot h^2 \cdot E / (m \cdot \Gamma^2)}$

m : not given E → Em, see above.

e: sqrt(1 + 2 · h^2 · Em / Γ^2);

$$\sqrt{1.258796152816846 \cdot 10^{-29} \cdot E_m \cdot h^2 + 1}$$

use : p1cf

q2: expand(pcf / m);

$$\frac{E}{m} = \frac{V^2}{2} - \frac{\Gamma}{r}$$

q2a: subst([E/m=Em, r=rh], q2);

$$E_m = \frac{V^2}{2} - 1.25 \cdot 10^{-7} \cdot \Gamma$$

q2b: ev(q2a, nouns);

$$E_m = -1.7825 \cdot 10^7$$

find : e

e: subst([q2b, sa], e);

$$0.307551394904985$$

6e) " conic section"

q2c: pe3b;

$$r = \frac{p}{e \cos(\theta - BB) + 1}$$

q2c1: ev(q2c, nouns);

$$r = \frac{1.0123345828934371 \cdot 10^7}{0.307551394904985 \cos(\theta - BB) + 1}$$

a) How to find : BB

BB is the phase angle =?, r=rh & θ = 0 (@ burnout)

q2c2: subst([r=rh, θ=0], q2c1);

$$8000000.0 = \frac{1.0123345828934371 \cdot 10^7}{0.307551394904985 \cos(BB) + 1}$$

q2c3: solve(q2c2, BB);

solve: using arc-trig functions to get a solution.

Some solutions will be lost. $\left[BB = \arccos\left(\frac{750250481621386347}{869347155820000000}\right) \right]$

q2c4: float(q2c3);

$$[BB = 0.529609391730455]$$

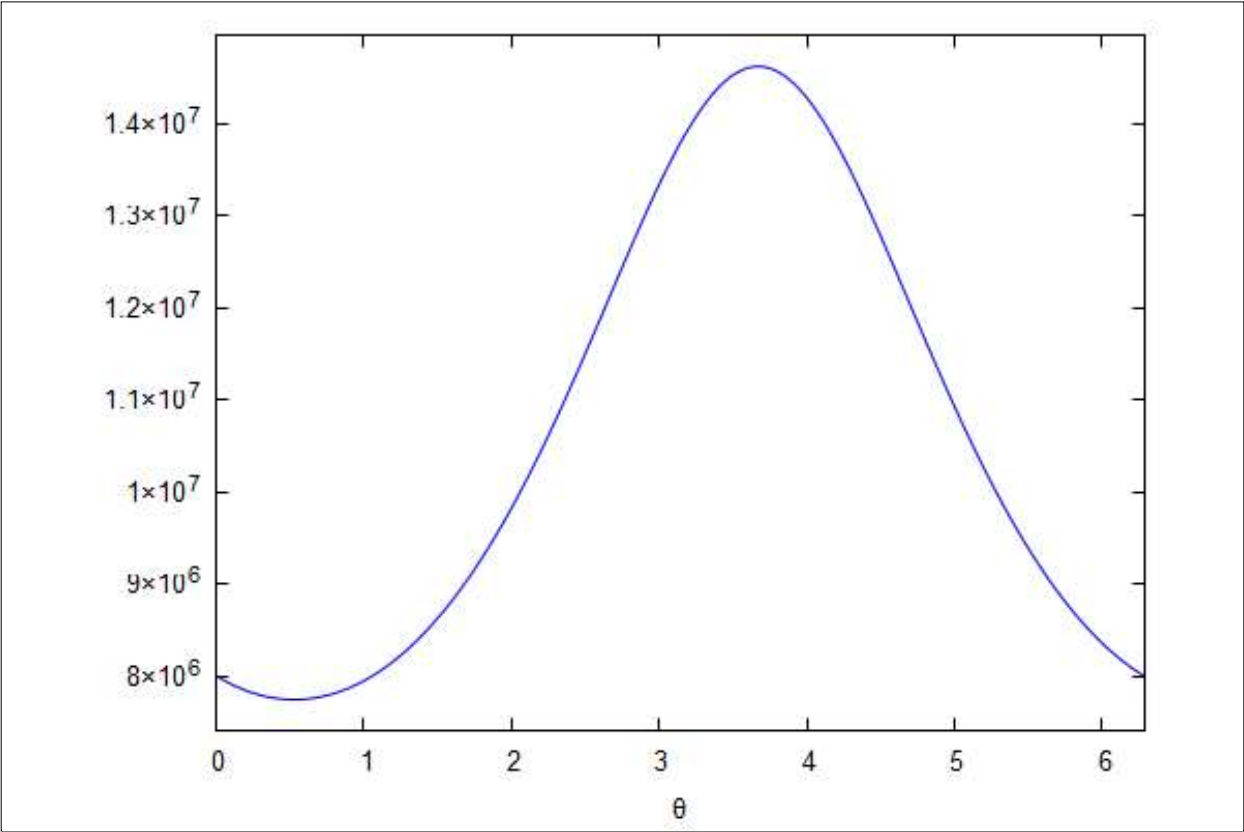
b) general equation : conic

use : q2c1

q2c5: subst(q2c4, q2c1);

$$r=\frac{1.0123345828934371\cdot 10^7}{0.307551394904985\cos(\theta-0.529609391730455)+1}$$

`wxplot2d(rhs(q2c5), [0,0,2*%pi])$`



because : value negative ,because cos(-x)=cos(x),automatic simplification.
